

#### **University of London**



National Centre for Earth Observation



### LEVERHULME

Centre for Wildfires, Environment and Society

# New AF Developments of European Satellites

Wooster, M., Xu, W., Nguyen, H., He, J., Roberts, G., Johnston, J., Strydome, T.,

## Structure



 Fire Radiative Energy eMissions (FREM) Products from Geostationary Satellites

 Sentinel-3 Polar Orbiting SLSTR Active Fire and FRP Products

### Fire Radiative Energy & Fuel Consumption



Fuel Consumption =  $\Sigma$  FRP  $\times$  CF

Theoretically ~ constant

National Centre for

### **GFAS MODIS-estimated FRE to GFED3.1 Fuel Consumption**



Kaiser et al. (2012) Biogeosciences

ALL

74%

0.85

## "Fire Radiative Energy eMissions" [FREM] Approach – "Fire Biome Map"





% tree cover > 5 m tall using 30 m Landsat 2015 Vegetation Continuous fields (VCF) product. FREMv2 African biome map using 300 m ESA CCI Landcover map (2015). Two woodland savanna biomes were separated into low and high % tree cover using the VCF product.

# **Smoke Plume Delineation**

DOY: 224 MODIS Slots: AQUA 2011-08-12T1240 224 V ihe2



SEVIRI Active Fire Detections MODIS AF Detections

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Link FRE released by a fire from first detection to measurement of plume AOD.

Datetime: Fri, 2011 August 12

FRE assessed from 15-min temporal resolution SEVIRI FRP-PIXEL Products



# **Smoke Plume Delineation**





# New Developments in FREMv2





# **Coefficient Derivation**





### Fuel Consumption Per m<sup>2</sup> of Burned Area



### **Extension to use Trace Gas Observations from S5P**

Emissions are derived from coefficients linking Meteosat-derived Fire Radiative Energy (FRE) totals to atmospheric species (CO) for different biomes - Nguyen et al. (*ACPD*)



Meteosat Active Fire pixels from which FRE is derived

### **Meteosat / S5P Fire Emissions Product Evaluation**

Emissions validation using emissions + WRF-CMAQ compared to independent Sentinel-5P CO observations



 111 23

Date

AUG 03

AUG 17

AUG 10

AUG 31

Any other species, x, can be estimated via emission factor ratios:

$$C_{e}^{x}\left[g.MJ^{-1}\right] = \frac{EF_{x}\left[g.kg^{-1}\right]}{EF_{reference}\left[g.kg^{-1}\right]} \cdot C_{e}^{reference}\left[g.MJ^{-1}\right]$$

Where reference species is CO

## Sentinel-3 Sea and Land Surface Temperature Radiometer (SLSTR)

SLSTR	Band center	Bandwidth	Albedo range	Reference albedo	SNR at reference
Channel	(μm)	(µm)	(%)	(%)	
<b>S1</b>	0.555	0.020	0 - 100	0.5 – 30	25 – 570
<b>S2</b>	0.659	0.020	0 - 100	0.5 – 30	25 – 570
<b>S3</b>	0.865	0.020	0 - 100	0.5 – 30	21 – 630
<b>S4</b>	1.375	0.015	0 - 100	0.5 – 5	25 – 162
S5	1.610	0.060	0 - 100	0.5 – 50	37 – 900
<b>S6</b>	2.250	0.050	0 - 100	0.5 – 4	27 – 142
			BT range (K)	Reference BT (K)	NEdT at reference (mK)
<b>S7</b>	3.74	0.38	200 – 311 🔭	270	56
<b>S8</b>	10.85	0.9	200 - 321	270 <sup>6</sup> 7 270	29
<b>S9</b>	12	1.0	200 – 318	270	21
F1	3.74	0.38	285 – 450	285 – 450	680 - 16
F2	10.85	0.9	230 - 400	230 – 400	79 – 35



- Two operational satellites
- Global Daily Data
- @ ~ Terra overpass times

# S3 SLSTR Data Intricacies





Footprint Overlap Near Nadir Scan

# Operational S3 AF Detection & FRP Products





- Global daily active fire (AF) counts and FRP retrievals from S3A and S3B.
- SLSTR has S7 and F1 MIR channels
- S7 often saturates by day (starting at BTs > 311 K) so day and night products are different

## S3 SLSTR AF Detection Alg. Structure



### Some Key Points

- Daytime synthesised "BT<sub>4</sub>" channel combing S7 & F1
- Background characterisation uses BT<sub>4</sub> < 330 K
- Contextual AF pixel detection thresholds increased compared to the night-time version.
- Sun-glint & desert detection thresholds further optimised.
- AF pixel clustering used to detect weak AF pixels at AF cluster edge



- Use S7 where possible for initial AF pixel detection.
- Always use F1 for FRP retrieval or not?

# F1\_ON and F1\_OFF Comparison





- Each fire has cluster of AF pixels in S7 and F1 (each imperfectly co-registered)
- F1\_OFF any fire without saturation S7 is processed using the S7 data (lower noise, more sensitive).
- F1\_ON means that such fires are anyway processed with F1 (smaller pixels)



## **Airborne Data Collection Over African Fires**



### Simulation of SLSTR F1\_ON & F1\_OFF Observations



290 300 310

320 >330 (K)

#### Always use F1 for 2<sup>nd</sup> Pass AF detection & FRP retrieval

### Global Nighttime SLSTR AF & FRP Example Comparison

**Active Fire Counts** 

**FRP Totals** 



Night-time S3 detects more AF pixels than MODIS, but less than VIIRS. FRP totals less different

## Night-time FRP Time-Series SLSTR vs. MODIS

South America



SLSTR S3B 150 125 125 125 100 5 10 15 20 25 30 Jan.2019

**MODIS** Terra

200



- Active fire pixel counts are higher for SLSTR than for MODIS.
- FRP values are far more similar.



### S3B FRP Product Performance Compared to MODIS



Per Pixel FRP Freq. Distribution (c)SLSTR MODIS 104 AF Pixel Count 103 102 20 0 40 100 120 140 AF Pixel FRP (MW)

### Global Daytime SLSTR AF & FRP Example Comparison





### Final S3 Spaceborne to Airborne FRP Dataset



### Retrieved FRP Uncertainty & Sub-Pixel Fire Position



- F1 channel shows benefits compared to S7 channel, despite higher noise.
- Even with F1\_ON, FRP can vary by up to ~ x2 depending on fire sub-pixel position.
- Comparison to airborne shows all but one fire to be within  $\sim x^2$ , and most far closer.
- F1\_ON confirmed as a far better option.
- F1\_ON now used in NTC products available from Sentinel Data Hub.

### Obtaining S3 AF Detection & FRP Products – Two Sources



[FIRMS Integration probably coming]

Opernicus Europe's eves on Earth Total number 1 km hot-spots = 2635 FRP 1 km: Total = 18793.0 [MW] - Avg. = 7.1±14.7 [MW] - Min = 0.2 [MW] - Max = 308.4 [MW] 70°N 50°N 30°N 10°N 10°S 4 30°S 50°S 70°S 90°S 160°W 120°W 80°W 0° 40°E 40°W 80°E 120°E 160°E 20 40 60 80 100 0

Sentinel-3 A SLSTR - Standard FRP MWIR [MW] - Night - 2.0 deg resolution - 19.06.2022

